

ALMA Offline Requirements - All Priorities (1, 2, & 3)

Test 1: AIPS++ Glish: single field interferometry, *only priority 1 requirements graded*

Test 2: AIPS++ Glish: small mosaics, *only priority 1 requirements graded*

Test 3: AIPS++ Glish: combining single-dish and mosaic synthesis images, *only priority 1 requirements graded*

Test 4: CASApY: Single BL interferometry, commissioning support, *only priority 1 requirements graded*

User Focus group: user interface working group input, *no grades*

Test 5: CASApY: 1) single dish reduction & imaging & 2) synthesis regression testing, *priority 1 & 2 requirements graded*

Grades: **A** = adequate, **E** = enhancement desired, **I** = inadequate, **U** = unable to evaluate

For grades with **(P)** attached, the requirement could only be partially tested.

Sect.	Req. No.	Priority	Requirement description	1st Audit Grade	Test 1 Grade	Test 2 Grade	Test 3 Grade	Test 4 Grade	Test 5 SD	Test 5 synth	Comments
1.1			General Rqmts & relation to the Pipeline								
1.1	R1	1	An offline data processing pkg must fulfill the rqmts in this document	A							
1.1	R2	1	All std obs modes supported by ALMA must be processable by the pkg.	U							
1.1	R3	1	Pkg shall be installable at users inst and available at ALMA centers	U	A	A	A		x	x	
1.1	R4	1	Pkg performance shall be benchmarked and meet ALMA rqmts	U			A(P)			x	
1.1	R5	2	Functionality available in pipeline should also be in the offline package	U							
1.2			Operational Issues								
1.2	R1	1	Installation of pkg shall be easy	A	A	A/E	A/E		x	x	
1.2	R2		Error reporting & handling shall be user-understandable & non-destructive:								
1.2	R2.1	1	Error reporting messages shall be written for end users, not programmers.	I	I	I	I		x	x	
1.2	R2.2	1	Must be provision for job ctrl (interrupt & abort)	A	I	I	I		x	x	
1.2	R2.3	2	Error handling shall be non-destructive (no corrupted data), fail gracefully	I					x	x	
1.2	R2.4	3	<i>Code traceback of execution errors shall be available</i>	A							
1.2	R3.1	1	Logging of commands and user inputs shall be provided	I	A/E	A/E	A/E		x	x	
1.2	R3.2	1	Logging of tool results shall be provided (errors, files written, timing)	A	A/E	A/E	A/E		x	x	
1.2	R3.3	2	Logging of output such as summaries of results shall be provided	A					x	x	
1.2	R3.4	2	The session log shall be readable by the user (e.g. text file, not binary)	A					x	x	
1.2	R3.5	3	<i>Session logs shall be executable by package UI, reproducing the session</i>	I							
1.2	R4	1	There shall be comprehensive handling of multi users/tasking	A					x		
1.2	R5	1	There shall be support for the pkg (bug fixes done according to priority)	U	A	A	A		x	x	
1.2	R6	1	Improvements & updates based on user feedback shall be done	A	A	A	A		x	x	
1.2	R7	2	Backward compatability of core functions should not be broken w.out reason	U						x	
1.2	R8	2	Source code for astronomical routines shall be available to the user	A							
1.2	R9	2	User installation of package not restricted by expensive or licenses	A						x	
1.2	R10	2	Provision for development and incorporation of user-supplied code possible.	A							
1.2	R11	2	Successive stages of call, correction, flagging should not destroy data.	A						x	
2.1			General User Interface Rqmts								
2.1	R1.1	1	A command line interface (CLI) shall be provided	A	A	A	A		x	x	
2.1	R1.2	1	A GUI shall be provided - GUI actions shall be scripted	A							
2.1	R2	1	The user shall be able to run host operating system commands (ls, rm)	A	A/E	A	A		x	x	
2.1	R3	1	Multi-tasking for all interfaces shall be available where appropriate	A					x	x	
2.2			Graphical User Interface (GUI)								
2.2	R1.1	1	GUI window updating must be fast (less than 0.1s)	U							
2.2	R1.2	1	Windows shall not take up excessive screen space (<=800x800 pixels)	I							
2.2	R1.3	2	Be able to choose cascading windows or re-use single window for new ops	I							
2.2	R1.4	2	Must be easy & robust to move and resize all windows	A							
2.2	R1.5	3	<i>There shall be a master control GUI for process control</i>	A/E							
2.2	R2	1	It must be easy to run GUIs remotely	A							
2.2	R3	2	GUIs should be easy to learn how to use	A/E							
2.2	R4	2	GUIs should be available for reduction of all standard ALMA obs modes	U							
2.2	R5	2	GUI look & feel should be uniform throughout the package	A							
2.2	R6		GUI features for beginner modes shall include:								
2.2	R6.1	2	built in help facility with access to novice-oriented help doc	A							

2.2	R6.2	2	Sensible defaulting of values for parameters, with guidance for user choice	A/E							
2.2	R6.3	2	Integrated functionally built around common data analysis tasks	I							
2.2	R7	3	<i>It shall be easy for users to develop & include custom GUIs in the package</i>	A							
2.3	R2		Command Line Interface (CLI)								
2.3	R1	1	Command files for batch CLI processing shall be possible	A	A	A	A	x	x		
2.3	R2	1	The CLI shall have command-line recall & editing w/ name completion	A/E	A/E	A/E	A/E	x	x		
2.3	R3	1	All GUI functionality must be available in CLI mode	A							
2.3	R4	2	CLI use easy over low-speed modem lines or network, with ASCII emulation	A							
2.3	R5	3	Desirable to have CLI mode menu to edit input parms	Missing				x	x		
2.4			Interface programming, parameter passing & feedback								
2.4	R1.1	1	The UI must have variable assignment & evaluation	A						x	
2.4	R1.2	1	The UI must have array handling	A						x	
2.4	R1.3	1	The UI must have conditional statements	A			A			x	
2.4	R1.4	1	The UI must have control loops	A			A			x	
2.4	R1.5	1	The UI must have string manipulation	A						x	
2.4	R1.6	1	The UI must have user-defined functions & procedures w/ arguments	A						x	
2.4	R1.7	1	The UI must have process control, interrupts, error handling	A						x	
2.4	R1.8	1	The UI must have std mathematical operations & functions	A						x	
2.4	R1.9	2	The UI must have efficient special vector and matrix operations	A							
2.4	R1.10	2	The UI must have user-defined data structures	A							
2.4	R2	1	Commands executed will be logged with provision to re-execute	A				x	x		
2.4	R3	1	Input parm. & syntax checking shall be done when a function is called	I				x	x		
2.4	R4	2	Parms should be passable between applications, global vars not the default	A				x	x		
2.4	R5	2	Vars shall be named consistently and clearly using astronomical terms	I				x	x		
2.4	R6		Parameter inputs to tools shall be stored for later recall:					x	x		
2.4	R6.1	2	Save inputs on closure and reinstate on next instance of tool	A				x	x		
2.4	R6.2	3	<i>Be able to save/recall state of parms for entire package & individual tools</i>	I							
2.5			Documentation & Help Facility								
2.5	R1.1	1	User cookbooks with extensive examples shall be available	I	A	A/E	A/E	x	x		
2.5	R1.2	1	Application help & ref manual (w/ all fctn & tool inputs) shall be available	I	A/E	A/E	A/E	x	x		
2.5	R1.3	1	On-line help, FAQs, e-mail contact info shall be available	A	A	A	A	x	x		
2.5	R1.4	1	Release history, bug reports/tracking, patch descriptions shall be available	I	A	A	A	x	x		
2.5	R1.5	2	Programmer references and guides shall be available	I							
2.5	R1.6	2	Data format descriptions shall be available	A				x	x		
2.5	R1.7	2	Algorithm descriptions shall be available	A/E				x	x		
2.5	R1.8	3	<i>news letters, e-mail exploders, notes series shall be available</i>	A							
2.5	R2	1	Documentation must be up-to-date & complete for all parts of the pkg.	I	I	I	I				
2.5	R3.1	1	Help mtl's shall be available in std document print format (pdf, postscript)	I	A/E	A/E	A/E	x	x		
2.5	R3.2	2	Help mtl's shall be available in printer-friendly versions of HTML pages	A				x	x		
2.5	R3.3	3	<i>Help mtl's shall be available in popular proprietary formats (MS-Word)</i>	Missing							
2.5	R4	2	Help shall be context-sensitive where relevant, includes GUI fly-over banners	A				x(P)	x(P)		evaluate only whether help is context-sensitive
2.5	R5	2	GUIs: help directs browser to web page. CLI: in-line text based help	A/E				x(P)	x(P)		Evaluate only CLI in-line text-based help
2.5	R6	2	Full search capability must be built into documentation library	A				x	x		
3.1			Data Handling - General Data Requirements								
3.1	R1		The pkg must support data taken in any std ALMA hdw modes	A							
3.1	R2.1	1	The pkg should handle program header info	A	A/E	A	A	x	x		
3.1	R2.2	1	The pkg should handle obs status info (and schedules)	A							
3.1	R2.3	1	The pkg should handle field info	A	A	A	A	x	x		
3.1	R2.4.1	1	The pkg should handle SD&Interf. data organized by position	A	A	A(P)	A(P)	x	x		
3.1	R2.4.2	1	The pkg should handle SD&Interf. data organized by subreflector state	A							
3.1	R2.4.3	1	The pkg should handle SD&Interf. data organized by polarization products	A	A	A(P)	A(P)	x	x		
3.1	R2.4.4	1	The pkg should handle SD&Interf. data organized by spectral channels	A	A	A(P)	A(P)	x	x		
3.1	R2.4.5	1	The pkg should handle SD&Interf. data organized by frequency bands	A	A/E	A(P)	A(P)	x	x		
3.1	R2.4.6	1	The pkg should handle SD&Interf. data organized by lfs	A	A	A(P)	A(P)	x	x		
3.1	R2.4.7	1	The pkg should handle SD&Interf. data organized by subarrays (incl. ACA)	A	A	A(P)					
3.1	R2.5.1	1	The pkg should handle cross-correlations (including 12m array x ACA)	A	A	A(P)	A(P)			x	
3.1	R2.5.2	1	The pkg should handle auto-correlations	A							
3.1	R2.5.3	1	The pkg should handle uncorrected &/or on-line WVR corrected data	A							
3.1	R2.5.4	1	The pkg should handle phase-array data (if chosen by user)	A							

3.1	R2.6	1	The pkg should handle total power from antennas in SD mode	A					x	
3.1	R2.7	1	The pkg should handle weights &/or data uncertainties	A	A/E	A	A		x	x
3.1	R2.8.1	1	The pkg should handle OTF scanning mode	A					x	
3.1	R2.8.2	1	The pkg should handle subreflector switching mode	A						
3.1	R2.8.3	1	The pkg should handle frequency switching mode	A						
3.1	R2.9	1	The pkg should handle flagging data or masks	A	A	A			x	x
3.1	R2.10	1	The pkg should handle diagnostic data & errors	A	I	I				
3.1	R2.11.1	1	The pkg should handle bandpass calibration data	A	A	A			x	x
3.1	R2.11.2	1	The pkg should handle source flux density cal data	A	A	A			x	x
3.1	R2.11.3	1	The pkg should handle antenna polarization leakage cal data	A						
3.1	R2.12.1	1	The pkg should handle gain tables	A	A	A				x
3.1	R2.12.2	1	The pkg should handle flux bootstraps	A	A/E	A/E			x	x
3.1	R2.13	1	The pkg should handle images &/or models produced from data	A	A	A	A		x	x
3.1	R2.14	1	The pkg should handle processing history	A	A	A	A		x	x
3.1	R3.1	1	The user should be able to select srcnames, etc (w/ wildcarding)	I					x	x
3.1	R3.2	1	The user should be able to select mosaic or scanning pointing centers	A		U			x	x
3.1	R3.3	1	The user should be able to select polarization products or channels	A	A	A	A		x	x
3.1	R3.4	1	The user should be able to select bands (e.g. frequency bands, lfs)	A	A	A	A		x	x
3.1	R3.5	1	The user should be able to select spectral channels	A	A/E	A/E	A		x	x
3.1	R3.6	1	The user should be able to select frequency-switched data	A						
3.1	R3.7	1	The user should be able to select interferometer subarrays	A						
3.1	R3.8	1	The user should be able to select interf. WVR-corrected/uncorrected BLs	A						
3.1	R3.9	1	The user should be able to select ACA subarrays	new						
3.1	R4	1	Multiple pointing centers for mosaics must be supported	A		A	A		x	x
3.1	R5	1	Image transformation to desired Stokes parameters must be provided	A						
3.1	R6	1	Averaging of data over time, bands and spectral channels shall be possible.	I	I	I	I		x	x
3.1	R7	1	Data taken in arbitrary scanning patterns must be dealt with	A						
3.1	R8.1	1	Universal Time (UT) shall be supported	A	A	A	A		x	x
3.1	R8.2	1	Coordinated Universal Time (UTC) shall be supported	A						
3.1	R8.3	1	International Atomic Time (IAT) shall be supported	A						
3.1	R8.4	1	Local sidereal time (LST) shall be supported	A						
3.1	R8.5	1	Greenwich Mean Sidereal Time (GMST) shall be supported	A						
3.1	R8.6	1	Julian date (JD) & Modified Julian Date (MJD) shall be supported	A						
3.1	R8.7	1	Dynamical Times (TDT, TDB) shall be supported	A						
3.1	R9.1	1	Equatorial coordinates (RA,Dec) shall be supported	A	A	A	A		x	x
3.1	R9.2	1	Ecliptic coords (ELON, ELAT) shall be supported	A						
3.1	R9.3	1	Heliocentric coords (GLON, GLAT) shall be supported	A						
3.1	R9.4	1	Galactic coords (GLON, GLAT) shall be supported	A						
3.1	R9.5	1	Supergalactic coords (SLON, SLAT) shall be supported	A						
3.1	R9.6	1	Terrestrial coords (Az, El) shall be supported	A						
3.1	R10.1	1	J2000 equinox (and other FK5 equinoxes) shall be supported	A	A	A	A		x	x
3.1	R10.2	1	B1950 (and other FK4 equinoxes) shall be supported	A	A	A	A		x	x
3.1	R10.3	1	Geocentric apparent place ref frame & coords of date shall be supported	A						
3.1	R10.4	1	Topocentric coords shall be supported	A						
3.1	R10.5	1	International celestial ref system (ICRS) shall be supported	A						
3.1	R11.1	1	Radio velocity definition shall be supported	A	A	A	A		x	x
3.1	R11.2	1	Optical velocity definition shall be supported	A						
3.1	R11.3	1	Redshift velocity definition shall be supported	A						
3.1	R12.1	1	Std velocity frames shall be supported, including: topocentric	A						x
3.1	R12.2	1	Std velocity frames shall be supported, including: geocentric	A						
3.1	R12.3	1	Std velocity frames shall be supported, including: barycentric	A						
3.1	R12.4	1	Std velocity frames shall be supported, including: heliocentric	A						
3.1	R12.5	1	Std velocity frames shall be supported, including: kinematic LSR	A						x
3.1	R12.6	1	Std velocity frames shall be supported, including: dynamic LSR	A						
3.1	R12.7	1	Std velocity frames shall be supported, including: galactocentric	A						
3.1	R12.8	1	Std velocity frames shall be supported, including: local group	Missing						
3.1	R13.1	1	Coords/locations defined w/ respect to topocentric frame shall be supported	A						
3.1	R13.2	1	Coords/locations defined w/ respect to geocentric frame shall be supported	A						
3.1	R14	1	Flagging mask/tables must be maintained with the data during all ops	A	A	A			x	x

3.3	R2.3	1	Elliptical Gaussian models must be supported	A							
3.3	R2.4	2	Uniform (optically thick) disk models must be supported	A							
3.3	R2.5	2	Optically thin disk models must be supported	Missing							
3.3	R2.6	3	<i>Wavelets must be supported</i>	Missing							
3.3	R2.7	3	<i>Pixons must be supported</i>	I							
3.3	R3	1	Pixel blanking shall be maintained through image processing	A							
3.3	R4		The package shall support all std projections, including:								
3.3	R4.1	2	- sine or slant orthographic (SIN)	A							
3.3	R4.2	2	- tangent or gnomonic (TAN)	A							
3.3	R4.3	2	- cylindrical plate caree (CAR)	A							
3.3	R4.4	2	- Mercator (MER)	A							
3.3	R4.5	2	- stereographic (STG)	A							
3.3	R4.6	2	- Hammer-Aitoff (AIT)	A							
3.4			Foreign Data								
3.4	R1	2	Be able to process data from other interferometers, SDs if in ALMA format	U				x	x		
3.4	R2	2	Be able to import FITS images and combine with ALMA data	U				x(P)	x(P)	test import of FITs images only	
3.5			Interaction with the Archive								
3.5	R1	1	Access from the archive must be supported	U							
3.5	R2	1	Pkg/archive interaction must not interfere w/ other access to the archive	U							
3.5	R3	1	Security/integrity of the archive must be ensured during access/interface	U							
4.1			General Calibration & Editing Requirements								
4.1	R1	1	The pkg shall be able to handle all ALMA std cal modes	U							
4.1	R2	1	Cal, editing, & data correction shall be easily reversible	A	A/E	A/E					
4.1	R2.1	2	Logging of editing steps shall be in history table or 'data object'	I				x	x		
4.1	R2.2	3	<i>Individual edit undo is desirable</i>	I						x	
4.1	R3	1	Cal solution intervals should be optimized for the data (e.g. scan-based sols)	Missing	A	A				x	
4.1	R4		Data editing shall be possible based on:	Missing							
4.1	R4.1	1	- ointing data	Missing							
4.1	R4.2	1	- array tracking info (encoders);	Missing							
4.1	R4.3	1	- weather data	I							
4.1	R4.4	1	- Tsys data	I							
4.1	R4.5	1	- WVR data	Missing							
4.1	R4.6	2	- RFI monitoring	Missing							
4.1	R4.7	2	- Site-test interferometer (STI) and/or tipping radiometer	Missing							
4.1	R4.8	3	- array monitoring points (e.g. dewar temps) if data provided in std format	Missing							
4.1	R4.9	3	- other site instruments (e.g. FTS) if data is provided in std format	Missing							
4.1	R5	1	Data average cal sols should be possible w/ controlled interpolation, etc.	I	I	I				x	
4.1	R6		Inderactive data editing, cal & display of cal quanties: graphical & intuitive								
4.1	R6.1		Editing display tools shall include specification by:								
4.1	R6.1.1	1	- antenna	A	A	A				x	
4.1	R6.1.2	1	- baseline	A	A	A				x	
4.1	R6.1.3	1	- timerange	A	A	A				x	
4.1	R6.1.4	1	- uvrange	A	A	A				x	
4.1	R6.1.5	1	- pointing center	I		U				x	
4.1	R6.1.6	1	- plot versus Az, El	Missing							
4.1	R6.1.7	1	- plot versus HA range	Missing							
4.1	R6.1.8	1	- plot versus parallactic angle	Missing							
4.1	R6.1.9	1	- plot a slice through a data cube	Missing							
4.1	R6.2	1	Data editing tools should display spectra and spectral cubes	I	A/E	A/E		x	x		
4.1	R6.3	1	Data editing tools should display closure quantities	Missing							
4.1	R6.4	1	Data editing tools should display & select monitor data (e.g. Tatm, Tamb)	Missing							
4.1*	R6.5	1	Editing tools shall plot all the above vs BL, time, band, chan averaging	(missing)	I	A/E				x	Evaluate for parms that are available
4.1	R6.6	1	Edt tools shall plot amp/ph. vs time, vs time-BL w/ zoom, select, & clipping	I	A	A/E				x	
4.1	R6.7		Editing of data shall be possible based on stastical quantities, including:								
4.1	R6.7.1	1	- a data point vs a running mean over a timescale	Missing	I	A/E				x	
4.1	R6.7.2	1	- a data point versus a median over a timescale	Missing	I	I				x	
4.1	R6.7.3	1	- an rms scatter in a time range	Missing	I	A/E				x	
4.1	R6.7.4	1	- difierence versus a model	Missing							
4.1	R6.8	1	Editing display tools shall have auto-scaling & user-specified axis scaling	A	A	A				x	

4.1	R6.9	1	Editing display tools shall have auto-scaling & user-spec color/greyscale	A	A	A				x	
4.1	R6.10	1	Editing display tools shall include/mark flagged data in plots & auto-scaling	A	A	A				x	
4.1	R7		Non-interactive/automatic editing tools shall be available, including								
4.1	R7.1	1	- ability to edit with all capabilities in interactive editing	A	A/E	A/E				x	
4.1	R7.2	2	- automated editing w/ tunable criteria for auto selection of parm ranges	A							
4.1	R8	2	Data cal. correction, flagging possible based on std or user models	Missing			A/E(P)				
4.1	R9	2	Access to time history of cal info shall be built into cal engines.	I						x	
4.1	R10	2	Be able to establish/verify relative cal between epochs, configs,data subsets	I							
4.1	R11	2	Data disply/editing tools shall be applicable to SD and synthesis data.	A					x	x	
4.1	R12	2	Editing shall be possible in all visualization tools (e.g. cal solns, other plots)	I							
4.2			Atmospheric Calibration								
4.2	R1	1	A std atm model for Pipeline data processing shall exist-User-supplied too	Missing							
4.2	R2		Predict atm absorption, emission, pathlength based on the following data:								
4.2	R2.1	1	- temp, pressure, humidity	Missing							
4.2	R2.2	1	- measured atm emission	Missing							
4.2	R2.3	2	- data from site test interferometer or tipping radiometer	Missing							
4.2	R2.4	3	- measured FTS data (if FTS available for ALMA)	U							
4.2	R2.5	3	- measured atm profiles of temp & water if available	U							
4.2	R3	1	Atm models shall derive Tsys corrected for atm absorption to correct amps	Missing							
4.2	R4	1	Atm models shall convert WVR data to astronomical phase in all bands.	Missing							
4.3			Interferometer data calibration								
4.3	R1.1	1	Default ant-based cal solutions shall be available for antenna gains	A	A	A				x	
4.3	R1.2	1	Default ant-based cal solutions shall be available for polarization leakages	A							
4.3	R1.3	1	Default ant-based cal solutions shall be avail for ant-dependent bandpasses	A	A	A				x	
4.3	R2.1	1	BL-based cal solutions shall be available for WVR corrections to some BLs	A							
4.3	R2.2	1	BL-based cal solutions shall be available for closure errors	Missing							
4.3	R2.3	1	BL-based cal solutions shall be available for BL-dependent bandpasses	Missing			A			x	
4.3	R3	1	Gain corrections based on models shall be possible w/ automatic editing	A/E	A/E	A/E				x	
4.3	R4	1	Cal quantities shall be transferable between srcs &/or freq bands	I	A/E	A				x	
4.3	R5	1	Determining time-variable cmplx bandpass & transfer to srcs shall be simple	I	I	I				x	
4.3	R6	1	Full polarization calibration shall be possible	A							
4.3	R7	2	Determining , correcting, examining, flagging closer errors should be easy	Missing							
4.3	R8	2	Using std models shall be easy for cal ops (user supplied models possible)	A							
4.3	R9	2	Use redundancy when possible to increase accuracy of or check cal solns.	Missing							
4.3	R10	3	Provide synth pointing, focus, baseline, beam response fitting in pkg	Missing							
4.4			Single Dish data calibration								
4.4	R1	1	Fitting of spectral bandpass from cal src obs should be available	A					x		
4.4	R1.1	3	Remove standing wave ripples by fitting sine function to line-free channels	new							
4.4	R1.2	3	Provide "robust" fitting functions which are more tolerant to RFI	new							
4.4	R2	1	De-striping/adjust must exist for SD OTF obs w/ overlap/crossing scans.	I					x		
4.4	R3	3	Support cal of system parms (temp-controlled loads, noise sources)	Missing							
4.4	R4	3	Be able to process pointing, focus, tipping, beam-fitting data for SD data	Missing							
4.4	R5	1	Be able to phase calibrate datasets with ACA 12m & 7m dishes	new							
4.4	R6		Be able to do sky subtraction with multiple techniques:								
4.4	R6.1	2	Specify an arbitrarily complex ref/source order for multiple sky subtractions	new							
4.4	R6.2	1	Support position switched sky subtractions	new					x		
4.4	R6.3	1	Support nutator switched sky subtraction	new							
4.4	R6.4	2	Support frequency switched sky subtractions	new					x		
4.4	R6.5	3	Handle basic Tsys, gain cal, opacity corrections in sky subtractions	new							
4.4	R7	1	Sky subtraction methods during slow OTF observing should be supported	new							
4.4	R8		All functions on SD data must support polarization:								
4.4	R8.1	2	Be able to calculate leakage terms from a cal source and correct the data	new							
4.4	R8.2	2	Be able to determine absolute P.A. from cal source & correct the data	new							
4.4	R9		Be able to easily flag individual spectra or channels:								
4.4	R9.1	1	User should be able to set whole or part of a spectra as invalid (flag)	new					x		
4.4	R9.2	1	User should be able to set individual or range of spectral points as invalid	new					x		
4.4	R9.2.1	2	User can flag a channel & interpolate value from adjacent good data	new							
4.4	R9.3	3	User can plot average spectral flux across band, interactively flag data	new							
4.4	R9.4	3	Fitting routines should be able to select RFI-tolerant (robust) algorithms	new							

4.4	R9.5	3	Automatically find birdies or RFI-corrupted data - indicate data as invalid	new							
4.5			Mosaicing Considerations for calibration								
4.5	R1	1	Determination of pting offsets & polarized primary beam must be available	U		A(P)				x(P)	evaluate correction for primary beam only
4.5	R2	2	Be able to cross check, correct relative cal between synth & SD mosaics	U							
4.5	R3	3	Allow user-input of scan pointing info from pattern or interpolated from list	A/E							
4.6			Ancillary & Diagnostic Data considerations for calibration								
4.6	R1	1	Output from atm monitoring (WVR, FTS) shall be importable to cal sftw	U							
4.6	R2	1	Cal solns/correction based on pting, focus, subreflector status shall exist	Missing							
4.6	R3	2	Be able to import for cal procedures environmental data (weather)	U							
4.6	R4	3	Be able to read and use in cal procedures engineering monitor data	U							
5.1			General Imaging Requirements								
5.1	R1	1	Imaging of all ALMA exported data must be provided	U							
5.1	R2	1	Efficient selection of subsets of imaging data must be provided	A	A/E	A/E	A/E			x	
5.1	R3		Provide a variety of imaging, deconvolution, analysis algorithms, including:								
5.1	R3.1	1	- 'dirty' images with selectable weighting	A	A	A	A			x	
5.1	R3.2	1	- residual images after model subtraction	A	A	A	A			x	
5.1	R3.3	1	- single-scale CLEAN deconvolution (Hogbom, Clark, Cotton-Schwab)	A	A	A	A			x	
5.1	R3.4	1	- maximum entropy deconvolution (MEM)	A	A	A	A			x	
5.1	R3.5	1	- linear mosaicing algorithms	A		U					
5.1	R3.6	1	- non-linear mosaicing algorithms	A		A	A			x	
5.1	R3.7	2	- non-negative least-squares (NNLS)	A/E							
5.1	R3.8	2	- multi-scale CLEAN	A/E						x	
5.1	R3.9	2	- model fitting (point, Gaussian, disk)	I						x(P)	evaluate model fitting with point models only
5.1	R3.10	2	- multi-frequency synthesis w/ different spectral models	I						x(P)	evaluate multi-freq synthesis, no spectral models
5.1	R3.11	3	- special function deconvolution (Pixon, wavelet)	I							
5.1	R4	1	Image pixel and spectral channel blanking must be supported	A							
5.1	R5	1	Interactive graphical selection of deconvolution region masks shall exist	A	A/E	A/E	A/E			x	
5.1	R6	2	Support multiple dataset input to tools (not previously concatenated)	I						x	
5.1	R7	2	Provide integrated deconvolution, sel-cal, editing tool for novices	A/E							
5.1	R8	2	Compare images on different coord systems, equinoxes, or projections	A						x	
5.1	R9	2	Transform & merge correctly image cubes using diff vel definitions/frames	A						x	
5.2			Interferometric Imaging								
5.2	R1	1	Incorporation of the polarized primary beam response of the array is required	I		A(P)	A(P)			x(P)	Use *unpolarized* primary beam
5.2	R2	1	Imaging of RR, LL, RL, LR or I,Q,U,V must be selectable and interchangeable	A/E	A	A	A			x	
5.2	R3	1	Imaging data from multiple epochs/configurations must be possible	A/E	A/E	A/E	A/E			x	
5.2	R4	1	Simultaneous multiple-field imaging & deconvolution must be supported	A		U	U			x	
5.2	R5	1	Cont. subtraction in uv & image planes shall be available & flexible	I	A/E	A				x	
5.2	R6	1	Include zero-spacing/short-spacing data (ALMA & non-ALMA) w/ weighting	I			A/E(P)			x(P)	Include zero-spacing data, non-ALMA
5.2	R7	2	There shall be the choice of FFT & DFT imaging (esp. for small datasets)	Missing							
5.2	R8	2	Be able to do near-field imaging of solar-system objects	A							
5.3			Mosaicing and Single Dish Imaging Considerations								
5.3	R1	1	Combine interferometer & single dish in mosaic imaging	A/E			A(P)				
5.3	R1.1	1	Cross correlation between 12 array & ACA must be supported	U							requirement text changed
5.3	R2	1	Incorporate (polarized) primary beam & pointing correction in mosaics.	I		A(P)					
5.3	R2.1	1	Primary beam calcs & corrections must take into account OTF effects	U		U					
5.3	R2.2	1	ALMA std beam images will be made available and distributed with the pkg	U							
5.3	R2.3	2	Be able to specify primary beam in analytic, tabular & ALMA provided format	A							
5.3	R3	1	Image & uv-plane gridding parms/interpolation shall be selectable	A/E		A	A				
5.3	R4	1	Provide scaling, de-stripping of scans on image-plane combo & SD OTF	A						x	priority changed
5.3	R4.1	1	Provide initial de-stripping using spatial baselines from non-emission regions	new						x	
5.3	R4.2	1	Define/implement algorithms to remove residual striping	new							
5.3	R5	2	Produce an image by combining data observed on different rasters	A							
5.3	R6	2	Apply pointing corrections to the data during imaging	Missing							
5.3	R7	2	Be able to mitigate the effects of non-coplanar baselines & sky curvature	A							
5.3	R8	1	Be able to grid SD data in a cube if desired	new						x	
5.3	R9	1	Automatically & manually adjust weights when combining SD & synth data	new							x
5.3	R9.1	1	Combine synth & SD data using image feathering technique	new							x
5.3	R9.2	1	Combine synth, SD data with joint deconvolution, manually set weights	new							x
5.3	R9.3	2	Combine synth, SD data with joint deconvolution, weights auto determined	new							x

7.3	R1.6	1	- parallactic angle;	Missing					A/E			
7.3	R1.7	1	- uv distance;	I	A/E	A			A/E		x	
7.3	R1.8	1	- u and/or v;	I	A/E	A			U		x	
7.3	R1.9	1	- baseline	A	A	A			A/E		x	
7.3	R1.10	1	- channel, freq, or vel	A	A	A			A/E	x	x	
7.3	R1.11	1	- azimuth and/or elevation;	Missing					A/E	x	x	
7.3	R2	1	select	I						x	x	
7.3	R3	3	<i>Be able to interpolate any tabulated val onto visibility of cal solution pt.</i>	I								
7.3	R1.1	1	Be able to plot ancillary data amps or single dish power vs AZ and EL	I						x		
7.4			Other ALMA Data									
7.4	R1		Be able to plot standard ALMA-format ancillary data, including:									
7.4	R1.1	1	- amplitude or SD power vs Az and El	I								
7.4	R1.2	2	- focus data and curves	U								
7.4	R1.3	2	- pointing data and offset vectors	U								
7.4	R1.4	3	- WVR output data	U								
7.4	R1.5	3	- holography and beam map data	U								
7.4	R1.6	3	- monitor point values (e.g. temperatures)	U								
7.5			Image Cube Manipulation									
7.5	R1	1	Histograms of pixel values must be produced for selected regions of cube.	I							x	
7.5	R2		It shall be possible to view subsets or slices of data cubes:									
7.5	R2.1	1	- for axes aligned with cube faces	A								
7.5	R2.2	3	- for arbitrarily aligned axes	Missing								
7.5	R3	1	Be able to plot a pixel value in diff. cube layers/images against each other	A								
7.5	R4	1	Data cubes must be viewable as movies with selectable rates & directions	A	A	A	A			x	x	
7.5	R5	2	Support interactive display of spectra corresponding to pixel or image region	A						x	x	
7.5	R6	2	Provide interactive display of 1D slice taken from 2D image	Missing								
7.5	R7	2	Be able to plot spectra on pseudo-grid - thumbnail spectra in panels	I						x		
7.6			Single dish plotting:	new								
7.6	R1	1	Line & histogram plots of spectra should be possible	new						x		
7.6	R2		The user should be able to specify:									
7.6	R2.1	1	- line thickness	new						x		
7.6	R2.2	1	- line style	new						x		
7.6	R2.3	1	- Character size	new						x		
7.6	R2.4	1	- colors	new						x		
7.6	R3	1	Be able to produce hard copies without an interactive plotter starting	new						x		
7.6	R4	3	Be able to flexibly select data to plot (any header values for any scans)	new								
7.6	R5	2	Be able to auto-average integrations of a scan for plotting.	new								
7.6	R6	2	Be able to step back & forth between many spectra in multi-page plots	new								
7.6	R7	2	With multi-panels, automatically update plot when dimensions change	new								
7.6	R8	2	Plot full and zoomed copy of spectrum on same plot (to see weak features)	new								
7.6	R9	2	Allow offset in Y direction when stacking multiple spectra on one plot	new								
7.6	R10	3	Plot auto update to reflect user processing (from CLI or GUI), turn on/off	new								
7.6	R11	3	Be able to plot individual ints in a "waterfall" plot with interactive zoom/pan	new								
7.6	R12	3	Be able to interactively edit waterfall plots	new								
7.6	R13	3	Be able to export waterfall plots to FITS images	new								
7.6	R14	1	Be able to plot line catalog overlays (individual catalogs)	new						x		
7.6		2	Be able to plot line catalog overlays (ALMA combined spectral line catalog)	new								
7.6	R14.1	2	Be able to plot user catalog overlays	new								
7.6	R14.2	3	Auto identify spectral lines if peak is above some threshold	new								
7.6	R15	1	Be able to plot residual data w/ or w/out subtraction of fit functions	new						x		
7.6	R16	2	Display basic header data on plot or next to it	new								
7.6	R16.1	3	Be able to define where the plot header info will appear on the plot	new								
7.6	R17	2	Be able to define simple annotations (text and simple graphics)	new								

8.1			Special Features: Simulations								
8.1	R1	2	Provide simulation capability - what ever pipeline has.								
	R1.1	1	Provide flexible simulation capability for the offline user too.	new						x	
8.1	R2	3	Provide single dish specific simulation capability.	new							
8.2			Special Features: VLBI								
8.2	R1	3	Be able to support VLBI processing of ALMA data	I							
8.3			Special Features: Solar System Object Observing								
8.3	R1		Be able to calculate and compensate for moving objects in solar system								
8.3	R1.1		Ephemerides must be provided for major solar system objects, including:								
8.3	R1.1.1	1	- sun	A							
8.3	R1.1.2	1	- moon	A							
8.3	R1.1.3	1	- planets	A							
8.3	R1.1.4	1	- major asteroids (all >50km dia.);	Missing							
8.3	R1.1.5	1	- known short period comets;	Missing							
8.3	R1.2	1	Ephemerides must be calculated in all available reference frames	A							
8.3	R1.3	2	Be able to import a user-supplied ephemeris in tabular form	A							
8.3	R1.4	2	Be able to calculate positions from user-provided orbital elements.	Missing							
8.3	R2	1	The pkg must carry out astronomy for moving sources	A							
8.3	R3	1	Compute visibility amp correction for dist from ephemeris or user-supplied	Missing							
8.3	R4		Calculate physical ephemeris quantities for solar system objects, including:								
8.3	R4.1	2	- sub-Earth latitude and longitude	Missing							
8.3	R4.2	2	- sub-solar latitude and longitude	Missing							
8.3	R4.3	2	- sub-solar angle of North Pole	Missing							
8.3	R4.4	2	- season	Missing							
8.3	R4.5	2	- phase angle	Missing							
8.3	R5		Do coord trans from sky plane to planetocentric projections in OL-3.3-R4 &:								
8.3	R5.1	2	- 3D backprojection (onto planetary sphere)	Missing							
8.3	R6	2	Include imaging corrections for the near-field	A							
8.3	R7		In addition to std models, provide the following:								
8.3	R7.1	2	- prolate and oblate ellipsoids	I							
8.3	R7.2	2	- limb-darkened disks (polynomial, cosn, Legendre polynomial)	Missing							
8.3	R8	3	Support 3D reconstruction of emission using obs of target at different aspect angles &/or rotational phases.	Missing							
8.4			Special Features: Pulsar Observing								
8.4	R1	3	Support time-series data folding into multiple phase bins for pulsar periods	Missing							
8.4	R2	3	Support at least 16 phase bins	A							
8.4	R3	3	Interpolate into phase bins from correlator output	Missing							
8.4	R4	3	Support SD - sky subtractions of pulsar observations	new							
8.5			Array Analysis Support								
8.5	R1		The package shall be able to calculate baseline delays (fringe fitting), e.g:								
8.5	R1.1	1	Phase versus frequency (slope tells you how good delays are)	new						U	
8.5	R1.2	1	Phase versus time (calculate rates)	new						U	